TOMOGRAPHY OF SMALL RESIDUAL STRESSES

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ABSTRACT. We study the inverse problem of determining the residual stress in Man’s model using tomographic data. Theoretically, the tomographic data are obtained at the zeroth approximation of geometrical optics for Man’s residual stress model. For compressional waves, the inverse problem is equivalent to the problem of inverting the longitudinal ray transform of a symmetric tensor field. For shear waves, the inverse problem, after the linearization, leads to another integral geometry operator which is called the mixed ray transform. Under some restrictions on coefficients, we are able to prove the uniqueness results in these two cases.

1. INTRODUCTION

Residual stresses are the stresses existing within an elastic body in the absence of external loads. They can be caused, for example, by heat treatment. The (Cauchy) residual stress tensor $\tilde{R}$ is symmetric, $\tilde{R}_{ij} = \tilde{R}_{ji}$, satisfies the equilibrium equation

$$\text{div} \tilde{R} = 0 \quad \text{in} \ M \quad (1.1)$$

and the traction-free boundary condition

$$\tilde{R} N = 0 \quad \text{on} \ \partial M, \quad (1.2)$$

where $M \subset \mathbb{R}^3$ is the domain occupied by an elastic body and $N$ is the normal to the boundary $\partial M$. Special technologies exist for creating approximately uniform compressive residual stresses inside materials since such stresses improve life and performance of the material, for example, toughened glass. On the other hand, tensile residual stresses or compressive residual stresses with big gradients can be detrimental to the life of a structural component. It is therefore desirable to design reliable methods to determine the residual stress in a body. Besides its practical value, the determination of residual stresses gives rise to many challenging mathematical questions as well. Based on different model equations for the residual stress and using different types of measurements, several versions of the inverse problem of unique determination of the residual stress were studied in [HU], [Ho], [IWY1], [IWY2], [IMN], [Ra] (see also [Ro1], [Ro2] for related results).

In this paper, we study the problem of recovering the residual stress tensor from the viewpoint of tensor tomography. To this end, we consider the time-stationary elastic system with residual stress that is assumed to be comparable with the inverse of the frequency. To be more precise, we set $\tilde{R} = R/\omega$, where $\omega \in \mathbb{R}$ is the (angular) frequency of an elastic wave. In what follows, we refer to $R$ as the residual stress tensor again although its physical dimension is time-stress = time-force/area. The time-stationary elastic wave equation is

$$\text{div} \sigma + \omega^2 \rho u = 0, \quad (1.3)$$

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where \( u \) is the displacement vector, \( \rho > 0 \) is the density and \( \sigma \) is the first Piola-Kirchhoff stress tensor. In view of Man’s model [Ma], the first Piola-Kirchhoff stress tensor is expressed through the displacement vector by the equation

\[
\sigma = \lambda (\text{tr} \, \varepsilon) I + 2\mu \varepsilon + \frac{1}{\omega} \left[ R + \nabla u \cdot R + \nu_1 (\text{tr} \, \varepsilon) (\text{tr} \, R) I + \nu_2 (\text{tr} \, R) \varepsilon + \nu_3 \left( (\text{tr} \, \varepsilon) R + (\text{tr} \, \varepsilon R) I \right) + \nu_4 (\varepsilon R + R \varepsilon) \right],
\]

(1.4)

where \( \varepsilon \) is the infinitesimal strain tensor, \( \lambda \) and \( \mu \) are Lamé’s parameters, and \( \nu_1, \ldots, \nu_4 \) are Man’s parameters. Ignoring the term with \( 1/\omega \) in (1.4), we obtain the standard Hooke’s law

\[
\sigma = \lambda (\text{tr} \, \varepsilon) I + 2\mu \varepsilon.
\]

(1.5)

We refer to (1.5) as the background isotropic medium for (1.4), while (1.4) is called the quasi-isotropic perturbation of the medium (1.5) caused by the residual stress tensor \( R \). Here, the term “quasi-isotropic” is used to emphasize that the perturbation smallness parameter \( 1/\omega \) coincides with the reciprocal of the frequency. The concept of the quasi-isotropic approximation or quasi-isotropic media was first introduced by Kravtsov [Kr] for the Maxwell equations (see also [KO]). This approximation is based on the method of geometrical optics with rays generated by the background isotropic medium. Its generalization to elastic waves is presented in [Sh, Ch. 7].

Similar to the method used in [Sh] for a general anisotropic perturbation of an isotropic elastic medium, we start with applying the classical ray method to (1.3) with the stress tensor given by (1.4). We restrict ourselves to considering the zeroth approximation of geometrical optics. As compared with the classical case of isotropic media, our formulas for the zeroth approximation have two distinct features. First, an additional factor appears in the formula for the amplitude of a compressional wave to describe the accumulation of the wave phase along a ray due to the residual stress. Second, the Rytov law for shear waves has got an additional term depending linearly on the residual stress. The inverse problem we study in this work is to determine the residual stress from the results of registration of the compressional or shear waves on the boundary of the medium under investigation. For compressional waves, the inverse problem is equivalent to the problem of inverting the longitudinal ray transform of a symmetric tensor field. For shear waves, the inverse problem, after the linearization, leads to another integral geometry operator that is called the mixed ray transform.

The paper is organized as follows. In Section 2, we discuss the quasi-isotropic approximation for Man’s model of residual stress. For compressional waves, we derive the formula for the amplitude; while for shear waves, we obtain the Rytov law. All the content of Section 2 is actually some modification of the corresponding arguments from [Sh, Section 7.1]. We have chosen the following compromise presentation style in Section 2: all details are presented for compressional waves while our arguments for shear waves are condensed as far as possible. In Section 3, we investigate the inverse problem of determining the residual stress from measurements of compressional waves on the boundary. The inverse problem of determining the residual stress using shear waves is discussed in Section 4.

2. QUASI-ISOTROPIC APPROXIMATION FOR RESIDUAL STRESSES

In this section, we derive the quasi-isotropic approximation for the residual stress model (1.4) in detail. We first write (1.4) in the form

\[
\sigma = \lambda (\text{tr} \, \varepsilon) I + 2\mu \varepsilon + \frac{1}{\omega} \left[ R + \nabla u \cdot R + \varepsilon \right],
\]

(2.1)
where \( c = c(R) \) is a fourth rank tensor depending linearly on \( R \). To write down the dependence explicitly, we reproduce (1.4) and (2.1) in Cartesian coordinates

\[
\sigma_{jk} = \lambda \varepsilon_{pp} \delta_{jk} + 2 \mu \varepsilon_{jk} + \frac{1}{\omega} \left( R_{jk} + \nabla_j u_p \cdot R_{pk} \right) \quad (2.2)
\]

and

\[
\sigma_{jk} = \lambda \varepsilon_{pp} \delta_{jk} + 2 \mu \varepsilon_{jk} + \frac{1}{\omega} \left( R_{jk} + \nabla_j u_p \cdot R_{pk} + c_{jklm} \varepsilon_{lm} \right). \quad (2.3)
\]

Comparing these two formulas, we deduce

\[
c_{jklm} = \nu_1 R_{pp} \varepsilon_{jk} \delta_{lm} + \frac{\nu_2}{2} R_{pp} (\delta_{jl} \delta_{km} + \delta_{jm} \delta_{kl}) + \nu_3 (R_{jk} \delta_{lm} + R_{lm} \delta_{jk}) + \frac{\nu_4}{2} (R_{jl} \delta_{km} + R_{km} \delta_{jl}). \quad (2.4)
\]

Since we are going to use curvilinear coordinates, we rewrite (2.4) and (2.3) in the covariant form

\[
c_{jklm} = \nu_1 (\text{tr}_E R) g_{jk} g_{lm} + \frac{\nu_2}{2} (\text{tr}_E R) (g_{jk} g_{km} + g_{jm} g_{kl}) + \nu_3 (R_{jk} g_{lm} + R_{lm} g_{jk}) + \frac{\nu_4}{2} (R_{jl} g_{km} + R_{km} g_{jl} + R_{kl} g_{jm} + R_{jm} g_{kl}), \quad (2.5)
\]

\[
\sigma_{jk} = \lambda \varepsilon_{pp} g_{jk} + 2 \mu \varepsilon_{jk} + \frac{1}{\omega} \left( R_{jk} + u_i^l \nu_3 R_{pk} + c_{jklm} \varepsilon_{lm} \right). \quad (2.6)
\]

Here \( (g_{jk}) \) is the Euclidean metric tensor such that \( |dx|^2 = g_{jk} dx^j dx^k \), \( (\text{tr}_E R) = g^{jk} R_{jk} \), \( \varepsilon_{lm} = g^{ij} g^{km} \varepsilon_{jk} \), and \( \varepsilon_{m} = g^{lm} \varepsilon_{lm} \). Recall that \( \varepsilon_{jm} = (u_i^l + u_i^m) / 2 \), \( u_i^l = g_{jk} u^k \),

\[
u_i^m = \frac{\partial u_i}{\partial x^m} - \Gamma^k_{lm} u_k \quad \text{and} \quad u_i^m = \frac{\partial u_i}{\partial x^m} + \Gamma^l_{km} u^k,
\]

where \( \Gamma^l_{km} \) are the Christoffel symbols given by

\[
\Gamma^l_{km} = \frac{1}{2} g^{lj} \left( \frac{\partial g_{kj}}{\partial x^m} + \frac{\partial g_{mj}}{\partial x^k} - \frac{\partial g_{km}}{\partial x^j} \right).
\]

In curvilinear coordinates, (1.3) is expressed as

\[
\sigma_{jk} \varepsilon^{k} + \omega^2 \rho u_j = 0, \quad (2.7)
\]

where \( \sigma_{jk} = g^{lm} \sigma_{jl} \) and

\[
\sigma_{jk} = \frac{\partial \sigma_{jk}}{\partial x^m} - \Gamma^l_{jm} \sigma_{lk} - \Gamma^l_{km} \sigma_{lj}.
\]

It is clear that the tensor \( c \) possesses the symmetries

\[
c_{jklm} = c_{kjlm} = c_{jmlk} = c_{lmjk} \quad (2.8)
\]

as follows from (2.5). The equilibrium equation (1.1) is now written as

\[
R_{jk} \varepsilon^k = 0. \quad (2.9)
\]

We consider propagation of small elastic waves in a medium described by (2.7) with the constitutive optics law (2.6). We reproduce some arguments of book [Sh]. The method of geometrical optics consists of representing a solution to the system by the asymptotic series

\[
u_j = e^{i \omega \tau} \sum_{m=0}^\infty \frac{u_j}{(i \omega)^m}, \quad \varepsilon_{jk} = e^{i \omega \tau} \sum_{m=-1}^\infty \frac{\varepsilon_{jk}}{(i \omega)^m}, \quad \sigma_{jk} = e^{i \omega \tau} \sum_{m=-1}^\infty \frac{\sigma_{jk}}{(i \omega)^m},
\]
where $\tau = \tau(x)$ is a real function (eikonal). We insert the series into the equations under consideration, implement differentiations and equate the coefficients at the same powers of the frequency $\omega$ on the left- and right-hand sides of the so-obtained equalities. In such a way we arrive at the infinite system of equations

$$\begin{align*}
    \varepsilon_{jk}^m &= \frac{1}{2} \left( u_{j;k}^m + u_{k;j}^m + u_{j}^{m+1} \tau_{k}^m + u_{k}^{m+1} \tau_{j}^m \right) \quad (m = -1, 0, \ldots), \\
    \sigma_{jk}^m &= \sigma_{jk}^{m+1} \tau_{k}^m - \rho \ u_{j}^{m+2} = 0 \quad (m = -2, -1, \ldots), \\
    \xi_{jk}^m &= \lambda \varepsilon_{jk}^m + 2 \mu \ v_{j}^m + \iota (\tau_{j}^m R^m_{k} u^p_{j} + R^m_{kp} u^{m-1} p_{j} + c_{jkpq} \varepsilon_{pq}^m) \quad (m = 0, 2, 3, \ldots), \\
    \frac{1}{\sigma_{jk}} &= \lambda \xi_{jk}^m + 2 \mu \ v_{j}^m + \iota (R_{jk} + \tau_{j}^m R^m_{k} u^p_{j} + R^m_{kp} u^{m-1} p_{j} + c_{jkpq} \varepsilon_{pq}^m),
\end{align*}$$

where it is assumed that $u = \varepsilon = \sigma = 0$. Observe that $\sigma$ is a symmetric tensor while $\sigma^m (m \geq 0)$ is not symmetric. Putting $m = -1$ in (2.10) and $m = -2$ in (2.11), we have

$$\begin{align*}
    \varepsilon_{jk}^m &= \frac{1}{2} \left( u_{j;k}^m + u_{k;j}^m \right) \\
    \varepsilon_{jk}^m &= \frac{1}{2} \left( u_{j;k}^m + u_{k;j}^m \right)
\end{align*}$$

and

$$\begin{align*}
    \sigma_{jk}^m &= \lambda \varepsilon_{jk}^m + 2 \mu \ v_{j}^m + \iota (\tau_{j}^m R^m_{k} u^p_{j} + R^m_{kp} u^{m-1} p_{j} + c_{jkpq} \varepsilon_{pq}^m) \quad (m = 0, 2, 3, \ldots), \\
    \frac{1}{\sigma_{jk}} &= \lambda \varepsilon_{jk}^m + 2 \mu \ v_{j}^m + \iota (R_{jk} + \tau_{j}^m R^m_{k} u^p_{j} + R^m_{kp} u^{m-1} p_{j} + c_{jkpq} \varepsilon_{pq}^m),
\end{align*}$$

Observe that the residual stress does not participate in (2.12), (2.15), and (2.16). This means that geometry of rays is the same for the background isotropic medium (1.5) and quasi-isotropic medium (1.4). As well known (see, for example, [Sh, p. 702]), (2.12) and (2.15)–(2.16) imply that $\tau$ satisfies the eikonal equation

$$|\nabla \tau|^2 = n^2,$$

where

$$n_2 = n_p^2 = \frac{\rho}{\lambda + 2 \mu} \quad \text{or} \quad n_2 = n_s^2 = \frac{\rho}{\mu}.$$

Next, we calculate the (complex) amplitude $A_p$ of a compressional wave by repeating arguments of [Sh, Section 7.1.3]. Recall that the amplitude $A_p$ is defined by

$$0 \ u = \frac{\lambda + 2 \mu}{\rho} (u, \nabla \tau) \nabla \tau = A_p |\nabla \tau|.$$

We fix a solution $\tau$ to the eikonal equation with $n = n_p$ and introduce ray coordinates in a neighborhood of a ray, i.e., curvilinear coordinates $x^1, x^2, x^3$ such that $x^3 = \tau$ and the coordinates surfaces $x^3 = x_0^3$ are orthogonal to the coordinates lines $x^1 = x_0^1, x^2 = x_0^2$ that are geodesics of the metric $ds^2 = n_2^2 g_{jk} dx^j dx^k$. In such coordinates

$$\begin{align*}
    -1 \sigma_{\alpha \beta} &= \lambda n g_{\alpha \beta} A_p, \\
    \sigma_{\alpha 3} &= 0, \\
    \sigma_{33} &= (\lambda + 2 \mu) n^{-1} A_p
\end{align*}$$

(see [Sh, formulas (7.1.23)]). Here Greek indices assume the values 1,2. Likewise, we can get that

$$\begin{align*}
    0 \xi_{\alpha \beta} &= \frac{1}{2} \left( u_{\alpha ; \beta}^0 + u_{\beta ; \alpha}^0 \right), \\
    0 \xi_{33} &= u_{3 ; 3}^0 + u_{3}^0, \\
    0 \sigma_{33} &= \lambda n^{-2} g_{\alpha \beta}^0 \xi_{\alpha \beta}^0 + (\lambda + 2 \mu) \xi_{33}^0 + i (n R_{33} A_p + c_{33pq} \varepsilon_{pq}^0) \\
    -1 \sigma_{3k}^0 &= n^2 \sigma_{33} - \rho \ u_3 = 0.
\end{align*}$$
Formulas (7.1.27)-(7.1.29) of [Sh] remain unchanged, i.e.,
\[ 0 \partial u_{\alpha; \beta} = \frac{1}{2} n \partial g_{\alpha \beta} A_p, \quad 0 \partial u_{3; 3} = n - 1 \partial A_p \partial \tau, \]
\[ -1 \sigma_{3 \alpha; \beta} = \mu n \partial g_{\alpha \beta} A_p, \]
\[ -1 \sigma_{33; 3} = (\lambda + 2 \mu) n - 1 \partial A_p \partial \tau + \left[ n - 1 \partial (\lambda + 2 \mu) \partial \tau + (\lambda + 2 \mu) n - 2 \partial n \partial \tau \right] A_p, \]
\[ -1 \sigma_{3k; k} = (\lambda + 2 \mu) n - 1 \partial A_p \partial \tau + \left[ \mu n g_{\alpha \beta} \partial g_{\alpha \beta} + n \partial (\lambda + 2 \mu) \partial \tau + (\lambda + 2 \mu) \partial n \partial \tau \right] A_p, \]
\[ 0 \varepsilon_{\alpha \beta} = \frac{1}{2} n \partial g_{\alpha \beta} A_p, \quad 0 \varepsilon_{33} = n - 1 \partial A_p \partial \tau + u_3. \quad (2.21) \]

Using (2.23) in (2.19) implies
\[ 0 \partial n^3 = (\lambda + 2 \mu) n \partial A_p \partial \tau + \frac{1}{2} \lambda n - 1 g_{\alpha \beta} A_p + (\lambda + 2 \mu) u_3 + i(nR_{33} A_p + c_{33pq} \varepsilon_{pq}). \]

Inserting this value for \( n^3 \) and value (2.22) for \( -1 \sigma_{3k; k} \) into (2.20), we arrive at the relation
\[ 2(\lambda + 2 \mu) n \partial A_p \partial \tau + \left[ \frac{1}{2} (\lambda + 2 \mu) n g_{\alpha \beta} A_p + n \partial (\lambda + 2 \mu) \partial \tau + (\lambda + 2 \mu) \partial n \partial \tau \right] A_p \]
\[ + [\lambda + 2 \mu) n^2 - \rho] u_3 + i(n R_{33} A_p + n^2 c_{33pq} \varepsilon_{pq}) = 0. \]

In view of \( n^2 = \rho / (\lambda + 2 \mu) \), the coefficient at \( u_3 \) in this formula is equal to zero. Thus, inserting into the last formula the expressions
\[ -1 \varepsilon_{\alpha \beta} = -1 \varepsilon_{\alpha \beta} = 0, \quad -1 \varepsilon_{33} = n^3 A_p, \]
that follows from (7.1.11), (5.1.20) and (7.1.22) of [Sh], we arrive at the equation for the amplitude \( A_p \)
\[ \partial A_p \partial \tau + \left( \frac{g_{\alpha \beta} \partial g_{\alpha \beta}}{2} + \frac{1}{2(\lambda + 2 \mu)} \frac{\partial A_p \partial \tau}{\partial n} + \frac{1}{2n} \frac{\partial n \partial A_p \partial \tau}{\partial \tau} + \frac{i n^2}{\lambda + 2 \mu} R_{33} + \frac{i n^4}{2(\lambda + 2 \mu)} c_{3333} \right) A_p = 0. \]

Using notation (5.1.22) and formula (5.1.23) of [Sh], we write this in the form
\[ \partial A_p \partial \tau \left[ \ln \left( A_p [1/2 n^{1/2}(\lambda + 2 \mu)]^{1/2} \right) \right] = -i \frac{n^2}{\lambda + 2 \mu} R_{33} - i \frac{n^4}{2(\lambda + 2 \mu)} c_{3333}, \]
which, together with \( n^2 = n^2 \), gives
\[ -i \frac{\partial}{\partial \tau} \left[ \ln \left( \frac{A_p}{\sqrt{J \rho v_p}} \right) \right] = -i \frac{R_{33}}{\rho v_p^2} - i \frac{n^4}{2 \rho v_p^6} c_{3333}, \]
where \( J \) is the geometrical spreading (see [Sh, (5.1.22)]) and \( v_p = 1/n_p \) is the velocity of compressional waves. This implies the following analogous of [Sh, formula (7.1.32)]:
\[ A_p = \frac{C}{\sqrt{J \rho v_p}} \exp \left[ -i \int \gamma \frac{1}{\rho v_p} R_{3j} \gamma^j \gamma^k \partial \tau - i \int \gamma \frac{1}{2 \rho v_p^6} c_{jkm} \gamma^j \gamma^k \gamma^m \partial \tau \right], \quad (2.24) \]
where \( \gamma^j = d \gamma^j / d \tau \) and \( C \) is a constant for a given ray \( \gamma \).

Next, we express the second integrand through \( R \). Since \( \gamma(\tau) \) is a geodesic of the Riemannian metric \( |d\gamma|^2 = v_p^2 |d\gamma|^2_R \), it satisfies
\[ g_{jk} \gamma^j \gamma^k = |\gamma|^2_R = v_p^2 |\gamma|^2 = v_p^2. \]
under the assumption that $\gamma$ is parameterized by the $\tau$-length, i.e., $|\dot{\gamma}|^2 = 1$. Taking this into account, we obtain from (2.5)

$$c_{jklm} \dot{\gamma}^{j} \dot{\gamma}^{k} \dot{\gamma}^{l} \dot{\gamma}^{m} = v_p^2 \left(2(\nu_3 + \nu_4) R_{jk} \dot{\gamma}^{j} \dot{\gamma}^{k} + v_p^2 (\nu_1 + \nu_2) \text{tr}_E R \right). \tag{2.25}$$

Here $\text{tr}_E R$ is the trace of $R$ with respect to the Euclidean metric $(g_{jk})$, i.e., $\text{tr}_E R = g^{jk} R_{jk}$. Since $\gamma$ is a geodesic of the Riemannian metric $h_{jk} = v_p^{-2} g_{jk}$, it is more natural to use the Riemannian trace $\text{tr} R = h^{jk} R_{jk} = v_p^2 \text{tr}_E R$. Thus, (2.25) takes the form

$$c_{jklm} \dot{\gamma}^{j} \dot{\gamma}^{k} \dot{\gamma}^{l} \dot{\gamma}^{m} = v_p^2 \left(2(\nu_3 + \nu_4) R_{jk} \dot{\gamma}^{j} \dot{\gamma}^{k} + (\nu_1 + \nu_2) \text{tr} R \right). \tag{2.26}$$

Substituting this value into (2.24), we obtain the final formula for the amplitude of the compressional wave

$$A_p = \frac{C}{\sqrt{J \rho v_p}} \exp \left[-i \int_{\gamma} \frac{1}{\rho v_p^4} \left((1 + \nu_3 + \nu_4) R_{jk} \dot{\gamma}^{j} \dot{\gamma}^{k} + \frac{1}{2} (\nu_1 + \nu_2) \text{tr} R \right) \ d\tau \right]. \tag{2.27}$$

The physical meaning of the formula is as follows: the residual stress $R$ distorts the phase of the compressional wave.

Finally, we consider propagation of shear waves by repeating arguments of Sections 7.1.4 and 7.1.5 of [Sh]. It can be easily checked that the residual stress $R$ does not change equations (7.1.43) of [Sh]. Therefore the same formula $A_s = C/\sqrt{J \rho v_s}$ is valid for the amplitude of the shear wave as for the background isotropic medium where $v_s = 1/n_s$ is the velocity of shear waves. Moreover, Rytov’s law in [Sh, formula (7.1.51)] does not change either. We reproduce the formula here

$$\left( \frac{D \eta}{d\tau} \right)_j = -i \frac{1}{4 \rho v_s^6} (\delta_j^l - \dot{\gamma}_j \dot{\gamma}^l) c_{jklm} \dot{\gamma}^k \dot{\gamma}^m \eta^l, \tag{2.28}$$

where $\eta$ is the polarization vector defined by $0_u^0 = A_s n_s^{-1} \eta_j$. Here $D/d\tau = \dot{\gamma}^j \nabla_j$ with the covariant derivative taken with respect to the Riemannian metric $d\tau^2 = v_s^{-2} |dx|^2_E = \hat{h}_{jk} dx^j dx^k$ and indices are raised with the help of the same metric, i.e., $\dot{\gamma}^k = \hat{h}^{jk} \dot{\gamma}_j$ and $\eta^l = \hat{h}^{jl} \eta_j$.

3. THE INVERSE PROBLEM FOR COMPRESSIONAL WAVES

First of all we emphasize the following feature of our approach. While considering the forward problem, we impose no boundary condition on the displacement vector $u$ at the boundary of the domain under consideration. Thus, we treat the problems as if the waves propagate in an unbounded medium, and use the boundary only as a surface at which the sources and detectors of oscillations are disposed. In fact, due to the reflection effects on the boundary, the possibility of registration of information that is used below as the data for inverse problems seems to be rather problematic. Here we will not settle this question but only attract reader’s attention to the fact of its existence.

In contrast to the content of the last paragraph, the following remark gives the possibility of measuring the data for compressional waves regardless to any boundary condition. There exists an alternative version of the geometrical optics method which is based on the analysis of propagation of the wave front of a non-stationary elastic wave, see [CJ]. In this version, the integral participating on (2.27) appears as the first order perturbation for the travel time of a compressional wave. More precisely, for a fixed geodesic $\gamma$ of the metric $|dx|^2_p = v_p^{-2} |dx|_E^2$ between two boundary points, let $T_{1/\omega}(\gamma)$ be the propagation
time of the compressional wave along \( \gamma \) in the quasi-isotropic medium (1.4) and \( T_0(\gamma) \) be the corresponding travel time in the background isotropic medium (1.5). Then
\[
T_{1/\omega}(\gamma) - T_0(\gamma) = -\frac{1}{2\omega} D(\gamma) + o(\frac{1}{\omega}),
\]
where
\[
D(\gamma) = \int_{\gamma} \frac{1}{\rho v_p^2} ((1 + \nu_3 + \nu_4) R_{jk} \dot{\gamma}^j \dot{\gamma}^k + \frac{1}{2} (\nu_1 + \nu_2) \text{tr} R) d\tau
\]
is just the integral participating in (2.27). The corresponding result is obtained in [CJ] in the case of a perturbation of the form
\[
\sigma_{jk} = \lambda \varepsilon_{pp} \delta_{jk} + 2\mu \varepsilon_{jk} + \frac{1}{\omega} \varepsilon_{jklm} \varepsilon_{lm}.
\]
By repeating arguments of [CJ], one easily sees that the result is true if the last formula is replaced with (2.3). Thus, data (3.1) can be obtained by measuring travel times for compressional waves. Let us consider an elastic wave initiated by a \( \delta \)-kind source at the initial point of \( \gamma \) which starts at the time \( t = 0 \). The wave will be a mixture of different body and surface waves including secondary waves caused by reflections at the boundary. Nevertheless, \( T_{1/\omega}(\gamma) \) is the first arrival time to the final point of \( \gamma \) since compressional waves are the fastest elastic waves. In the simplest case of constant parameters \( \lambda, \mu, \rho \) \( T_0(\gamma) \) is equal, up to a constant factor, to the Euclidean distance between the endpoints of the straightline segment \( \gamma \). So, the only problem is the sufficiently precise measurement of the travel time \( T_{1/\omega}(\gamma) \).

Studying the inverse problem, we will first consider the case of constant coefficients \( \lambda, \mu, \rho, \nu_1, \ldots, \nu_4 \) since this case is much easier than the general one and, most probably, this case is of the most importance for applications.

For the inverse problems studied here, we assume the material parameters \( \lambda, \mu, \rho, \nu_1, \ldots, \nu_4 \) to be given a priori and only the residual stress \( R \) to be unknown. In practice, some of these parameters are also unknowns to be determined. Therefore the inverse problem of recovering residual stresses, when some of material parameters \( \lambda, \mu, \rho, \nu_1, \ldots, \nu_4 \) are also unknowns to be recovered, is also worth of investigation. But this new inverse problem is much harder because it is a \textit{nonlinear} problem.

3.1. The case of constant coefficients. Let the medium under consideration be contained in a bounded convex domain \( M \subset \mathbb{R}^3 \) with smooth boundary \( \partial M \) and let each of the material parameters \( \lambda, \mu, \rho, \nu_1, \ldots, \nu_4 \) be constant. In this case, the Riemannian metric \( h = v_p^{-2} |dx|^2_E \) is a constant multiple of the Euclidean metric \( g = |dx|^2_E \) and geodesics are intersections of straight lines with \( M \). The equilibrium equation (2.9) means that \( R \) is a solenoidal tensor field. We extend \( R \) to the whole of \( \mathbb{R}^3 \) by zero outside \( M \). Then the extended tensor field \( R \) is solenoidal on the whole of \( \mathbb{R}^3 \) in virtue of the boundary condition (1.2).

We study the inverse problem of recovering the residual stress tensor field. To this end, assume that we can dispose a source of compressional waves at every point of the boundary \( \partial M \) and measure the phase of a compressional wave on the same surface \( \partial M \). In virtue of (2.27), our data are integrals (3.1) that are known for every line \( \gamma \) of \( \mathbb{R}^3 \). Initially in (3.1), \( \gamma \) is parameterized by the arc length in the metric \( h \), i.e., \( |\gamma|_E = v_p = \text{const} \) and the trace is understood with respect to the metric \( h \), i.e., \( \text{tr} R = v_p^2 \text{tr}_E R \). After a simple rescaling, we obtain the same formula (3.1), where now \( |\gamma|_E = 1 \) and \( \text{tr} R \) is replaced by the Euclidean trace \( \text{tr}_E R \). From now on in this subsection, we use the Euclidean metric only and write \( \text{tr} R \) instead of \( \text{tr}_E R \).
To get a well defined inverse problem, we have to impose some restrictions on the material parameters. Indeed, if for example \( \nu_1 + \nu_2 = 2(1 + \nu_3 + \nu_4)/3 \), then the integrand on (3.1) is identically equal to zero for \( R = g \). Therefore, in this section, we assume that

\[
3(\nu_1 + \nu_2) + 2(1 + \nu_3 + \nu_4) \neq 0.
\]  (3.2)

The first term of the integrand on (3.1) is considered as the leading term. Therefore, in this section, we also assume that

\[
1 + \nu_3 + \nu_4 \neq 0.
\]  (3.3)

Introducing the notations

\[
f = 1 + \nu_3 + \nu_4 R, \quad a = \frac{\nu_1 + \nu_2}{2(1 + \nu_3 + \nu_4)}
\]  (3.4)

\( f \) is zero outside \( M \), we write data (3.1) as

\[
(I(f + a(\text{tr} f)g))(\gamma) = \int_\gamma (f_{jk} + a(\text{tr} f)g_{jk}) \gamma^j \gamma^k d\tau,
\]  (3.5)

where \( I \) is the (longitudinal) ray transform on \( \mathbb{R}^3 \) which is defined in Section 2.1 of [Sh].

Let us remind the theorem on decomposition of a tensor field into solenoidal and potential parts (Theorem 2.6.3 of [Sh]): every symmetric tensor field \( u = (u_{jk}) \in L^2 \) on \( \mathbb{R}^3 \) can be uniquely represented as

\[
u = \tilde{u} + dv, \quad \delta \tilde{u} = 0,
\]

where \( \nu \) is continuous outside \( M \) and satisfies \( \nu(x) \to 0 \) as \( |x| \to \infty \). Here the operators \( d \) (inner derivative) and \( \delta \) (divergence) are defined in local coordinates by formulas \( (dv)_{jk} = (\nabla_j v_k + \nabla_k v_j)/2 \) and \( (\delta u)_j = g^{kl} \nabla_k u_{jl} \) respectively, \( \nabla \) being the covariant derivative with respect to the Euclidean metric \( g \). The summands \( \tilde{u} \) and \( dv \) of the decomposition are called the solenoidal and potential parts of the tensor field \( u \) respectively.

We now investigate the question of uniqueness of a solution to the inverse problem. Let a solenoidal field \( f \) satisfy \( I(f + a(\text{tr} f)g) = 0, \) \( a = \text{const.} \) By Theorem 2.15.1 of [Sh], this means that the solenoidal part of \( f + a(\text{tr} f)g \) is equal to zero. Since the solenoidal part of \( f \) coincides with \( f \), we obtain the equation

\[
f + aS((\text{tr} f)g) = 0,
\]  (3.6)

where \( S(u) \) denotes the solenoidal part of a tensor field \( u \). Recall that the Fourier transform interweaves the operators \( S \) and \( T \), where \( Tu \) stands for the tangential part of \( u \), see Section 2.6 of [Sh] for details. Applying the Fourier transform to the last equation, we obtain

\[
\hat{f} + a(\text{tr} \hat{f})\varepsilon = 0,
\]  (3.7)

where the tensor field \( \varepsilon = T(g) \) is expressed in Cartesian coordinates by \( \varepsilon_{jk}(y) = \delta_{jk} - y_j y_k/|y|^2 \), \( y \) being the variable in the Fourier space.

Applying the operator \( \text{tr} \) to equation (3.7), we obtain \( (1 + 2a)\text{tr} \hat{f} = 0 \). Thus, the inequality \( 1 + 2a \neq 0 \) is the necessary and sufficient condition for the uniqueness of a solution to the inverse problem. Recalling (3.4), we write the condition as

\[
\nu_1 + \nu_2 + \nu_3 + \nu_4 \neq -1.
\]  (3.8)

If (3.8) holds, equation \( (1 + 2a)\text{tr} \hat{f} = 0 \) gives \( \text{tr} \hat{f} = 0 \). Then (3.7) implies \( \hat{f} = 0 \) and therefore \( f = 0 \).


Under hypothesis (3.8), an explicit inversion formula for recovering a solenoidal tensor field \( f \) from the data \( I(f + a(\text{tr} f)g) \) can be easily derived from the corresponding inversion formula for \( I \), see Theorem 2.12.2 of [Sh]. The corresponding stability estimate can be also obtained on the base of the Plancherel formula for the ray transform, see Section 2.15 of [Sh]. Moreover, to recover a solenoidal \( f \), we do not need to measure ray integrals \( (I(f + a(\text{tr} f)g))(\gamma) \) for all lines \( \gamma \) of \( \mathbb{R}^3 \). Repeating arguments of [Sh2], we see that three families of lines are sufficient for an effective reconstruction algorithm, each family consists of all lines parallel to a coordinate plane.

If (3.8) does not hold, i.e., if \( \nu_1 + \nu_2 + \nu_3 + \nu_4 = -1 \), then the space of solenoidal tensor fields \( f \) satisfying \( I(f + a(\text{tr} f)g) = 0 \) can be explicitly described. Indeed, in this case \( \text{tr} f \) can be an arbitrary function and equation (3.6) gives \( f = S(\alpha g) \) with an arbitrary scalar function \( \alpha \).

3.2. The case of variable coefficients. Let again \( M \subset \mathbb{R}^3 \) be a closed bounded domain with smooth boundary \( \partial M \). Now the material parameters \( \lambda, \mu, \rho, \nu_1, \ldots, \nu_4 \) are assumed to be known smooth functions of a point \( x \in M \). Let \( v_p = \sqrt{(\lambda + 2\mu)/\rho} \) be the velocity of compressional waves. By \( g \) we denote the Euclidean metric and by \( h = v_p^{-2}g \), the Riemannian metric corresponding to compressional waves. Assume \((M, h)\) to be a convex non-trapping manifold (CNTM) in the sense of definition given in [Sh1]. The same definition is presented in Section 4.1 of [Sh] but the term “compact dissipative Riemannian manifold” is used instead of CNTM.

We consider the inverse problem of recovering the residual stress tensor field. Our data are integrals (3.1) that are known for every geodesic \( \gamma \) of the metric \( h \) with endpoints in \( \partial M \). The geodesic is parameterized by the arc length in the metric \( h \), i.e., \( |\dot{\gamma}| = 1 \). By the same arguments as above, we assume inequalities (3.2) and (3.3) to be valid everywhere in \( M \). On using the same notations (3.4), we write data (3.1) as

\[
I(f + a(\text{tr} f)h) = \int_{\gamma} (f_{jk} + a(\text{tr} f)h_{jk}) \dot{\gamma}^j \dot{\gamma}^k \, d\tau, \tag{3.9}
\]

where \( I \) is the (longitudinal) ray transform on the CNTM \((M, h)\) which is defined in Section 4.2 of [Sh].

For a compact Riemannian manifold \((M, h)\), the theorem on decomposition of a tensor field into solenoidal and potential parts (Theorem 3.3.2 of [Sh]) is valid in the following form: every symmetric tensor field \( u = (u_{jk}) \) can be uniquely represented as

\[
u = \tilde{u} + dv, \quad \delta \tilde{u} = 0, \quad v \big|_{\partial M} = 0
\]

where the operators \( d \) (inner derivative) and \( \delta \) (divergence) are defined in local coordinates by the same formulas \((dv)_{jk} = (\nabla_{\gamma}v_k + \nabla_kv_j)/2\) and \((\delta u)_j = h^{kl}\nabla_ku_{lj}\) respectively, \( \nabla \) being the covariant derivative with respect to the metric \( h \).

According to Theorem 4.3.3 of [Sh], the solenoidal part of the tensor field \( f + a(\text{tr} f)h \) can be uniquely recovered from data (3.9) under the assumption

\[
k^+(M, h) < 1/3, \tag{3.10}\]

where \( k^+(M, h) \) is some curvature characteristic of the CNTM \((M, h)\). For the metric \( h = v_p^{-2}g \), condition (3.10) holds if the function \( v_p \) is sufficiently \( C^2 \)-close to a constant, the degree of the closeness depends on the size of the domain \( M \). For such a manifold, the null-space of \( I \) consists of potential fields that can be represented in the form \( dv \) with a covector field \( v \) satisfying the boundary condition \( v \big|_{\partial M} = 0 \).
For variable coefficients, the main difficulty relates to the following circumstance: the equilibrium condition (2.9) does not mean anymore that $f$ is a solenoidal tensor field with respect to the metric $h$. Indeed, (2.9) can be rewritten in terms of $f$ as

$$\delta_E(bf) = 0 \quad \text{with} \quad b = \frac{\rho v^4}{1 + \nu_3 + \nu_4},$$

(3.11)

where $\delta_E$ is the divergence with respect to the Euclidean metric $g$. We are going to prove, at least under some restrictions on the material parameters $\lambda, \mu, \rho, \nu_1, \ldots, \nu_4$, that a tensor field $f$ is uniquely determined by data (3.9) if it satisfies (3.11).

First of all we will rewrite equation (3.11) in terms of the divergence $\delta f$ with respect to the metric $h$. Denote $c = \nu_2^2$, then $g = ch$. On using standard formulas of tensor analysis, one easily calculates

$$\delta_E(bf) = bc^{-1}(\delta f) + c^{-1}\left(f_{jk}\nabla^k b - \frac{1}{2}bc^{-1}(\text{tr } f)\nabla_j c - f_{jk}\nabla^k c\right),$$

where the covariant derivative and trace are understood with respect to the metric $h$. Therefore (3.11) is equivalent to the equation

$$\delta f + b^{-1}f_{jk}\nabla^k b + \frac{1}{2}c^{-1}f_{jk}\nabla^k c - \frac{1}{2}c^{-1}(\text{tr } f)\nabla_j c = 0.$$

Denoting

$$\alpha = b^{-1}\nabla b + \frac{1}{2}c^{-1}\nabla c, \quad \beta = -\frac{1}{2}c^{-1}\nabla c,$$

(3.12)

we write the equation in the coordinate free form

$$\delta f + f\alpha + (\text{tr } f)\beta = 0.$$  

(3.13)

We now investigate the uniqueness question to the inverse problem. Let a tensor field $f$ be such that $I(f + a(\text{tr } f)h) = 0$. We assume that the Riemannian manifold $(M, h)$ is a CNTM and satisfies the curvature condition (3.10). Then, by Theorem 4.3.3 of [Sh], $f + a(\text{tr } f)h$ must be a potential field, i.e., there exists a covector field $v$ on $M$ satisfying the boundary condition $v|_{\partial M} = 0$ such that

$$f + a(\text{tr } f)h = dv.$$

Taking the trace of both parts, we obtain

$$(1 + 3a)(\text{tr } f) = \delta v.$$

By (3.2), $1 + 3a$ does not vanish on $M$ and we can write

$$\text{tr } f = (1 + 3a)^{-1}\delta v, \quad f = dv - a(1 + 3a)^{-1}(\delta v)h.$$

Substituting these values into (3.13), we arrive at the boundary value problem on the covector field $v$

$$\begin{cases} 
(\frac{a}{1 + 3a}d\delta - \delta d)v - (dv)\alpha + \frac{1}{1 + 3a}(\delta v)\left(\alpha - \beta + (1 + 3a)\nabla(\frac{a}{1 + 3a})\right) = 0, \\
v|_{\partial M} = 0.
\end{cases}$$

(3.14)

We have thus proved

**Proposition 3.1.** Let a three-dimensional CNTM $(M, h)$ satisfy the curvature condition (3.10) and let $a, \alpha$, and $\beta$ be defined by (3.4) and (3.12). Every symmetric tensor field $f$ satisfying (3.13) can be uniquely recovered from data (3.9) if and only if the boundary value problem (3.14) has no nontrivial solution.
Equation (3.14) is rather complicated in the case of general coefficients $a, \alpha$, and $\beta$. We are going to investigate the boundary value problem in the case when the background medium is sufficiently close to a homogeneous one, i.e., when $a$ is close to a constant and $\alpha$ and $\beta$ are small. Even in this case, we need to impose some restrictions on $a$.

Let us first find a condition that guarantees ellipticity of the boundary value problem.

**Proposition 3.2.** The operator \( \frac{a}{1+3a} \delta d - \delta d \) is elliptic if and only if the inequality
\[
|1 + a| < |1 + 3a|
\]
holds on the whole of $M$.

**Proof.** Recall that the principal symbols of the operators $d$ and $\delta$ are $\sqrt{-1}i\xi$ and $\sqrt{-1}j\xi$ respectively, where $i\xi$ is the symmetric multiplication by the covector $\xi$ and $j\xi$ is the contraction with $\xi$, see [Sh, Section 3.3] for details. Therefore the principal symbol of our operator is
\[
\sigma \left( \frac{a}{1+3a} \delta d - \delta d \right) = \left( j\xi i\xi - \frac{a}{1+3a} i\xi j\xi \right).
\]
By [Sh, Lemma 3.3.3], $j\xi i\xi = \frac{1}{2}(|\xi|^2 E + i\xi j\xi)$ on covectors, where $E$ is the identity operator. Therefore
\[
\sigma \left( \frac{a}{1+3a} \delta d - \delta d \right) = \frac{1}{2}(|\xi|^2 E + \kappa i\xi j\xi) \quad \text{with} \quad \kappa = \frac{1}{1+3a}.
\]
On the other hand, for a covector $v$,
\[
\langle (|\xi|^2 E + \kappa i\xi j\xi)v, v \rangle = |\xi|^2 |v|^2 + \kappa \langle \xi, v \rangle^2.
\]
Hence, (3.15) guarantees the positiveness of $|\xi|^2 E + \kappa i\xi j\xi$ for all $\xi \neq 0$. \(\square\)

Under suitable assumptions on coefficients, the triviality of a solution to the boundary value problem (3.14) is guaranteed by the following

**Theorem 3.3.** Given a three-dimensional CNTM $(M, h)$, let $D$ be its diameter, i.e., the length of the longest geodesic. Assume the coefficients $a, \alpha$, and $\beta$ of equation (3.14) to satisfy
\[
3a_0 + \frac{1}{2} a_0 + \frac{3}{2} \beta_0 + \frac{1}{4} (a_0^3 + \beta_0^3) D^2 < 1,
\]
where
\[
a_0 = \sup \left| \frac{a}{1+3a} \right|, \quad a_0 = \left( \sup |a| \right)^{1/2}, \quad \beta_0 = \left( \sup \left| \frac{a\alpha - \beta}{1+3a} \right| \right)^{1/2}.
\]
Then the boundary value problem (3.14) has only trivial solution.

The values of $a_0$ and $\beta_0$ can be made arbitrary small by assuming the background medium to be sufficiently close to a homogeneous one. Therefore the main part of hypothesis (3.17) is $a_0 < 1/3$ that is equivalent to
\[
a > -1/6
\]
It is interesting to compare (3.18) with the ellipticity condition (3.15). If $-1/6 < a < 0$, then the boundary value problem is not elliptic but has only trivial solution for sufficiently small $\alpha$ and $\beta$. If both conditions (3.15) and (3.17) are satisfied, we can use the standard stability estimate for the elliptic boundary value problem with the trivial kernel to obtain some stability estimate in the inverse problem of recovering a tensor field $f$ from data (3.9). This stability estimate will be similar to that of [Sh, Theorem 4.3.4]. We do not present it here.

To prove Theorem 3.3 we need the following
Lemma 3.4. For a covector field \( v \) on a three-dimensional CNTM \( (M, h) \) satisfying \( v|_{\partial M} = 0 \), the estimate
\[
\|v\|_{L^2}^2 \leq \frac{D^2}{10} (2\|dv\|_{L^2}^2 + \|\delta v\|_{L^2}^2)
\]
holds where \( D \) is the diameter of \( (M, h) \).

Proof. Let \( \Omega M = \{(x, \xi) \mid x \in M, \xi \in T_x M, |\xi|^2 = h_{ij}\xi^i\xi^j = 1\} \) be the unit sphere bundle. Introduce the \( L^2 \)-norm
\[
\|u\|^2_{L^2(\Omega M)} = \frac{1}{4\pi} \int_{\Omega M} |u(x, \xi)|^2 \, d\omega_x(\xi) dV(x),
\]
where \( d\omega_x \) is the volume form on the sphere \( \Omega_x = \Omega M \cap T_x M \) and \( dV \) is the Riemannian volume form on \( M \).

Given a covector field \( v \) on \( M \), define two functions on \( \Omega M \)
\[
\varphi(x, \xi) = v_i(x)\xi^i, \quad \psi(x, \xi) = (dv(x))_{ij}\xi^i\xi^j.
\]
The functions are related by the equation \( H\varphi = \psi \) where \( H \) is the differentiation with respect to the geodesic flow, see [Sh, Section 4.4] for details. If \( v|_{\partial M} = 0 \), then \( \varphi \) vanishes on the boundary of \( \Omega M \). Applying the Poincaré inequality [Sh, Lemma 4.5.1] with the weight \( \lambda \equiv 1 \), we obtain
\[
\|\varphi\|^2_{L^2(\Omega M)} \leq \frac{D^2}{2} \|H\varphi\|^2_{L^2(\Omega M)} = \frac{D^2}{2} \|\psi\|^2_{L^2(\Omega M)}. \tag{3.19}
\]

One can easily see that (compare with [Sh, Lemma 7.4.2]) in the three-dimensional case
\[
\|\varphi\|^2_{L^2(\Omega M)} = \frac{1}{3} \|v\|^2_{L^2(M)},
\]
\[
\|\psi\|^2_{L^2(\Omega M)} = \frac{1}{15} (2\|dv\|^2_{L^2(M)} + \|jdv\|^2_{L^2(M)}) = \frac{1}{15} (2\|dv\|^2_{L^2(M)} + \|\delta v\|^2_{L^2(M)}).
\]
Inserting these values into (3.19), we get the statement of the lemma. \( \square \)

Proof of Theorem 3.3. Let \( v \) be a solution to the boundary value problem (3.14). First of all we rewrite equation (3.14) in the form
\[
d\left( \frac{a}{1 + 3a} \delta v \right) - \delta dv - (dv)\alpha + (\delta v) \frac{a\alpha - \beta}{1 + 3a} = 0.
\]
Take the \( L^2 \)-product of the equation with \( v \), use the boundary condition \( v|_{\partial M} = 0 \) and the fact that \( d \) and \( -\delta \) are dual operators (see [Sh, Theorem 3.3.1])
\[
\|dv\|^2 = \left( \frac{a}{1 + 3a} \delta v, \delta v \right) + ((dv)\alpha, v) - \left( \frac{a\alpha - \beta}{1 + 3a} \delta v, v \right). \tag{3.20}
\]
We estimate each summand on the right-hand side of (3.20). Obviously
\[
\left| \left( \frac{a}{1 + 3a} \delta v, \delta v \right) \right| \leq a_0 \|\delta v\|^2. \tag{3.21}
\]
The second summand on the right-hand side of (3.20) is estimated as follows:
\[
|((dv)\alpha, v)| \leq (\alpha_0^{1/2}\|dv\|)(\alpha_0^{3/2}\|v\|) \leq \frac{1}{2} (\alpha_0\|dv\|^2 + \alpha_0^3\|v\|^2).
\]
This implies with the help of Lemma 3.4
\[
|((dv)\alpha, v)| \leq \frac{5\alpha_0 + \alpha_0^3D^2}{10} \|dv\|^2 + \frac{\alpha_0^3D^2}{20} \|\delta v\|^2. \tag{3.22}
\]
The last term on the right-hand side of (3.20) is estimated in the same way:

$$\left| \frac{\alpha c - \beta}{1 + 3\alpha} \delta v, \delta v \right| \leq \frac{\beta_0^2 D^2}{10} \| \delta v \|^2 + \frac{10\beta_0 + \beta_0^3 D^2}{20} \| \delta v \|^2. \tag{3.23}$$

With the help of (3.21)–(3.23), (3.20) gives

$$\| \delta v \|^2 \leq \frac{1}{10} (5\alpha_0 + \alpha_0^3 D^2 + 3\beta_0 D^2) \| \delta v \|^2 + \left( \alpha_0 + \frac{1}{20} (10\beta_0 + \alpha_0^3 D^2 + 3\beta_0 D^2) \right) \| \delta v \|^2. \tag{3.24}$$

The inequality $$\| \delta v \|^2 \leq 3 \| \delta v \|^2$$ is obvious since $$\delta v$$ is the trace of the $$3 \times 3$$-matrix $$(\delta v)_{ij}$$. With the help of the latter inequality, (3.24) gives

$$\left[ 1 - 3\alpha_0 - \frac{1}{2} \alpha_0 - \frac{3}{2} \beta_0 - \frac{1}{4} (\alpha_0^3 + 3\beta_0^3) D^2 \right] \| \delta v \|^2 \leq 0.$$ 

The coefficient in the brackets is positive under hypothesis (3.17) and the inequality implies $$\delta v \equiv 0$$. With the help of the boundary condition $$v|_{\partial M} = 0$$, this implies $$v \equiv 0$$. $\square$

4. THE INVERSE PROBLEM FOR SHEAR WAVES

Let $$v_s = \sqrt{\mu/\rho}$$ be the velocity of shear waves. By $$g$$ we denote the Euclidean metric and by $$h = v_s^{-2} g$$, the Riemannian metric corresponding to shear waves. Assume $$(M, h)$$ to be a CNTM. We assume that, for every geodesic $$\gamma : [0, 1] \to M$$, $$|\gamma|^2 = h_{jk} \gamma^j \gamma^k = 1$$ with endpoints in $$\partial M$$, we can activate the shear wave with an arbitrary initial polarization $$\eta(0)$$ and can measure the final polarization $$\eta(l)$$. Mathematically, this means that the fundamental matrix $$U(\gamma)$$ of system (2.28) is known for every geodesic $$\gamma$$ with endpoints in $$\partial M$$ such that $$\eta(l) = U(\gamma)\eta(0)$$. We are going to recover the residual stress $$R$$ from the data $$U(\gamma)$$.

Introduce the tensor field $$f$$ by

$$f_{jklm} = -i \frac{1}{4\rho v_s^6} (c_{jklm} + c_{jkm}) \tag{4.1}$$

It possesses the symmetries

$$f_{jklm} = f_{kjlm} = f_{jkml} = f_{lmjk}$$

as follows from (2.8). Express $$f$$ through $$R$$ by substituting (2.5) into (4.1)

$$f_{jklm} = -i \frac{1}{4\rho v_s^6} \left( \nu_1 (\text{tr} R) (h_{jlh_{km}} + h_{jmh_{kl}}) + \nu_2 (\text{tr} R) (2h_{jk}h_{lm} + h_{jlh_{km}} + h_{jmh_{kl}}) + \nu_3 (R_{jlh_{km}} + R_{jmh_{kl}} + R_{kjlh_{jm}}) + \nu_4 (2R_{jk}h_{lm} + R_{jlh_{km}} + R_{jmh_{kl}} + R_{kjlh_{jm}} + R_{klh_{jm}} + 2R_{lmh_{jk}}) \right). \tag{4.2}$$

Here $$\text{tr} R = h^{jk} R_{jk}$$ is the Riemannian trace.

Rytov’s law (2.28) is written in terms of $$f$$ as

$$\left( \frac{D\eta}{d\tau} \right)_j = (\delta^p_j - \delta^q_j \delta^p) f_{qjk} \gamma^q \gamma^r \eta^k. \tag{4.3}$$

To write Rytov’s law in a coordinate free form, we introduce the following notations.

Define the linear operator $$f_\xi$$ by $$(f_\xi \eta)_j = f_{qjk} \gamma^q \gamma^r \eta^k$$. For a nonzero vector $$\xi$$, let $$P_\xi$$ be the orthogonal projection onto $$\xi^\perp = \{ \eta | \langle \xi, \eta \rangle = \xi^i \eta_i = 0 \}$$. In coordinates $$(P_\xi \eta)_j = (\delta^p_j - \frac{1}{|\xi|^2} \xi^p \xi^j) \eta_p$$. Then the coordinate free form of (4.3) is

$$\frac{D\eta}{d\tau} = P_\xi (f_\xi \eta).$$
This equation can be also written in the form
\[
\frac{D\eta}{d\tau} = (P_\gamma f_\gamma)\eta, \quad \eta(0) = \eta_0, \quad (4.4)
\]
where, for a symmetric tensor \( u = (u_{jk}) \),
\[
(P_\xi u)_{jk} = (\delta^p_k - \frac{1}{|\xi|^2}[\xi_j\xi^p](\delta^q_j - \frac{1}{|\xi|^2}[\xi_k\xi^q])u_{pq},
\]
compare with [Sh, formula (5.2.1)]. \( P_\xi \) is the orthogonal projection onto the subspace \{ \( u \mid u_{jk}\xi^k = 0 \) \}. Let \( \dot{\gamma}^+ (\tau) \) be the (complex) two-dimensional subspace of (the complexification of) the tangent space \( T_{\gamma(\tau)}M \) consisting of vectors orthogonal to \( \dot{\gamma}(\tau) \), and let \( \Pi^{0,\tau}_\gamma : \dot{\gamma}^+(0) \to \dot{\gamma}^+(\tau) \) be the parallel transport along \( \gamma \) with respect to the metric \( h \). The solution to the initial value problem (4.4) can be written as
\[
\eta(l) = \Pi^{0,l}_\gamma (U(\gamma)\eta_0) \quad \text{for} \quad \eta_0 \in \dot{\gamma}^+(0) \quad (4.5)
\]
with some linear operator (the fundamental matrix) \( U(\gamma) : \dot{\gamma}^+(0) \to \dot{\gamma}^+(0) \). Observe that \( U(\gamma) \) is a unitary operator since the matrix of system (4.4) is skew-Hermitian as is seen from (4.2). We consider the problem of recovering the tensor field \( R \) from the data \( U(\gamma) \) known for all geodesics with endpoints in \( \partial M \).

The problem is strongly nonlinear since the data \( U(\gamma) \) depends on \( R \) in a nonlinear manner. Let us linearize the problem. To this end we represent the fundamental matrix by the Neumann series
\[
U(\gamma) = E + \int \Pi_{\gamma,0}^{0,\tau}((P_\gamma f_\gamma)(\tau)) \, d\tau + \int \Pi_{\gamma,0}^{\tau,0}((P_\gamma f_\gamma)(\tau)) \, d\tau \int \Pi_{\gamma,0}^{\tau,\tau}((P_\gamma f_\gamma)(t)) \, dt + \ldots,
\]
where \( E \) is the identity matrix, and delete the terms that are nonlinear in \( f \). In other words, we use Born’s approximation
\[
U(\gamma) - E = \int \Pi_{\gamma,0}^{0,\tau}((P_\gamma f_\gamma)(\tau)) \, d\tau
\]
as the data for the linearized inverse problem. Since the integrand is a symmetric operator, the data are equivalent to the quadratic form
\[
\langle (U(\gamma) - E)\eta(0), \eta(0) \rangle = \int \langle \Pi_{\gamma,0}^{0,\tau}((P_\gamma f_\gamma)(\tau)) \eta(\tau), \eta(\tau) \rangle \, d\tau \quad (4.6)
\]
on the two-dimensional vector space of vector fields \( \eta(\tau) \) that are orthogonal to \( \dot{\gamma} \) and parallel along \( \gamma \) in the sense of the metric \( h \), i.e., satisfy \( \frac{D\eta}{d\tau} = 0 \). We denote this space by \( \gamma^+ \). Since
\[
\langle (P_\gamma f_\gamma)\eta, \eta \rangle = \langle P_\gamma (f_\gamma \eta), \eta \rangle = \langle f_\gamma \eta, P_\gamma \eta \rangle = \langle f_\eta, \eta \rangle = f_{jklm}\dot{\gamma}^j\dot{\gamma}^k\eta^l\eta^m,
\]
(4.6) can be written as
\[
\langle (U(\gamma) - E)\eta(0), \eta(0) \rangle = \langle (Lf)(\gamma), \eta \rangle := \int f_{jklm}(\gamma(\tau))\dot{\gamma}^j(\tau)\dot{\gamma}^k(\tau)\eta^l(\tau)\eta^m(\tau) \, d\tau. \quad (4.7)
\]
The operator \( L \) defined by this formula is called the mixed ray transform (compare with [Sh, formula (7.1.55)]).
We express the integrand of (4.7) through $R$ by substituting value (4.2) for $f$. On using the relations $|\gamma|=1$ and $\langle\gamma,\eta\rangle=0$, we obtain
\[
(Lf)(\gamma, \eta) = -i \int_\gamma \frac{1}{4\rho v_s^4} \left( \nu_4 R_{jk} \eta^j \eta^k + \nu_4 R_{jk} \dot{\gamma}^j \dot{\gamma}^k |\eta|^2 + \nu_2 (\text{tr} R) |\eta|^2 \right) d\tau. \tag{4.8}
\]
The first term of the integrand on (4.8) is considered as the leading term. Therefore we assume in this section that the function $\nu_4$ does not vanish in $M$. Note that we do not use the equilibrium condition (2.9) in this section. Introducing the notations
\[
F = -\frac{i\nu_4}{4\rho v_s^4} R, \quad a = \frac{\nu_2}{\nu_4},
\]
we write (4.8) as
\[
(Lf)(\gamma, \eta) = \int_\gamma F_{jk} \eta^j \eta^k d\tau + |\eta|^2 \int_\gamma F_{jk} \dot{\gamma}^j \dot{\gamma}^k d\tau + |\eta|^2 \int_\gamma a(\text{tr} F) d\tau. \tag{4.10}
\]
The first term on the right-hand side of (4.10) is the transverse ray transform $(JF)(\gamma, \eta)$, compare with [Sh, formula (5.1.72)]. The second term coincides, up to the factor $|\eta|^2$, with the longitudinal ray transform $(IF)(\gamma)$, while the last term coincides with $I(a(\text{tr} F)h)(\gamma)$.

Therefore (4.10) can be written as
\[
(JF)(\gamma, \eta) + |\eta|^2 (IF)(\gamma) + |\eta|^2 I(a(\text{tr} F)h)(\gamma) = (Lf)(\gamma, \eta) \quad (\eta \in \gamma^\perp). \tag{4.11}
\]
We consider (4.11) as an equation in an unknown symmetric tensor field $F = (F_{jk})$ while the right-hand side $(Lf)(\gamma, \eta)$ is given.

Equation (4.11) can be simplified. Indeed, let $(\eta_1, \eta_2)$ be an orthonormal basis of $\gamma^\perp$. Then
\[
(JF)(\gamma, \eta_1) + (JF)(\gamma, \eta_2) + 2(IF)(\gamma) + 2I(a(\text{tr} F)h)(\gamma) = (Lf)(\gamma, \eta_1) + (Lf)(\gamma, \eta_2). \tag{4.12}
\]
Since
\[
(JF)(\gamma, \eta_1) + (JF)(\gamma, \eta_2) = \int_\gamma (\text{tr} P_\gamma F) d\tau,
\]
(4.12) can be written as
\[
\frac{|\eta|^2}{2} \int_\gamma (\text{tr} P_\gamma F) d\tau + |\eta|^2 (IF)(\gamma) + |\eta|^2 I(a(\text{tr} F)h)(\gamma) = \frac{|\eta|^2}{2} \left( (Lf)(\gamma, \eta_1) + (Lf)(\gamma, \eta_2) \right).
\]
Subtracting this equality from (4.11), we obtain
\[
(JF)(\gamma, \eta) - \frac{|\eta|^2}{2} \int_\gamma (\text{tr} P_\gamma F) d\tau = (Lf)(\gamma, \eta) - \frac{|\eta|^2}{2} \left( (Lf)(\gamma, \eta_1) + (Lf)(\gamma, \eta_2) \right). \tag{4.13}
\]
Let $Q_\gamma$ be the orthogonal projection of symmetric tensors onto the subspace $\{u = (u_{jk}) | u_{jk} \xi^k = 0, \text{tr} u = 0\}$ (compare with Section 6.2 of [Sh]). The left-hand side of (4.13) can be transformed as follows:
\[
(JF)(\gamma, \eta) - \frac{|\eta|^2}{2} \int_\gamma (\text{tr} P_\gamma F) d\tau = \int_\gamma \left( (P_\gamma F) \eta, \eta \right) - \frac{|\eta|^2}{2} (\text{tr} P_\gamma F) \right) d\tau
\]
\[
= \int_\gamma ((Q_\gamma F) \eta, \eta) d\tau = (KF)(\gamma, \eta),
\]
where $K$ is the truncated transverse ray transform, see the definition in Section 6.2 of [Sh]. Since $(Lf)(\gamma, \eta)$ is known for every $\eta \in \gamma^\perp$, the right-hand side of (4.13) is known too. Moreover, it is independent of the choice of an orthonormal basis $(\eta_1, \eta_2)$ as follows from (4.13). Denoting the right-hand side of (4.13) by $D(\gamma, \eta)$, we arrive to the equation

\[ (KF)(\gamma, \eta) = D(\gamma, \eta) \quad (\eta \in \gamma^\perp). \tag{4.14} \]

Let us distinguish the trace free part of the tensor $F$, i.e., represent it in the form

\[ F = \tilde{F} + \frac{1}{3}(\text{tr } F) h, \quad \text{where} \quad \text{tr } \tilde{F} = 0. \]

By Theorem 6.6.2 of [Sh], the trace free tensor field $\tilde{F}$ can be uniquely recovered from $KF$ if the CNTM $(M, h)$ satisfies some curvature condition. Moreover, the stability estimate

\[ \| \tilde{F} \|_{L^2} \leq C \| KF \|_{H^1} \]

holds with some constant $C$ independent of $F$.

On assuming $\tilde{F}$ has been recovered, we can calculate $I\tilde{F}$ and $J\tilde{F}$. Then

\[
(JF)(\gamma, \eta) + |\eta|^2 (IF)(\gamma) = \frac{1}{3} (J((\text{tr } F) h))(\gamma, \eta) + |\eta|^2 (I((\text{tr } F) h))(\gamma, \eta) + (J\tilde{F})(\gamma, \eta) + \frac{1}{3} (I((\text{tr } F) h))(\gamma, \eta).
\]

Substituting this value into (4.11), we obtain the equation

\[
(I((a + \frac{2}{3}))(\text{tr } F))(\gamma) = \frac{1}{|\eta|^2} \left( (L f)(\gamma, \eta) - (J\tilde{F})(\gamma, \eta) \right) - (I\tilde{F})(\gamma). \tag{4.15}
\]

The left-hand side of this equation is the ray transform of the scalar function $(a + \frac{2}{3})(\text{tr } F)$ while the right-hand side is known. By the way, the right-hand side must be independent of $\eta \in \gamma^\perp$ as follows from the equation.

By Mukhometov’s theorem [Mu], the function $(a + \frac{2}{3})(\text{tr } F)$ can be uniquely recovered from the ray transform $I((a + \frac{2}{3})(\text{tr } F))$ if $(M, h)$ is a simple manifold, see Section 1.1 of [Sh] for the definition of a simple manifold. Thus, the trace $\text{tr } F$ can be recovered under the additional assumption that the function $3a + 2$ does not vanish. Since the trace free part $\tilde{F}$ has been already recovered, this gives the uniqueness statement for a solution to equation (4.11). The corresponding stability estimate can be also obtained.

Finally, let us discuss the case of constant material parameters $\lambda, \mu, \rho, \nu_1, \ldots, \nu_4$. In this case, to solve equation (4.11), we do not need to measure ray integrals $(Lf)(\gamma, \eta)$ for all lines $\gamma$ of $\mathbb{R}^3$. Indeed, as is proved in [LS], three families of lines are sufficient to recover the trace free part $\tilde{F}$ from the data $KF$, each family consists of all lines parallel to a coordinate plane. To solve equation (4.15), it suffices to know the right-hand side for all lines $\gamma$ parallel to a plane.

REFERENCES


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